# Lecture12: Quiz 2 Review Plasma Waves

**APPH E6101x** Columbia University



- Linearized equations of "fluid" motion
- Plane wave representation ("phasors")
- "Cold" plasma without pressure: describes the simplest plasma wave properties
- "Warm" plasma includes (usually) electron pressure: easiest for ion acoustic waves
- Electrostatic (longitudinal) and Electromagnetic waves
- Polarization (and wave energy density)
- Dispersion relations,  $\omega$  vs. k, phase-velocity, group-velocity
- Wave propagation through a plasma as a diagnostic

# Key Concepts

$$\mathbf{E} = \hat{\mathbf{E}} \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)]$$
$$\mathbf{B} = \hat{\mathbf{B}} \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)]$$
$$\mathbf{j} = \hat{\mathbf{j}} \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)]$$



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# **Basic Equations Review**

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$
$$\nabla \times \mathbf{B} = \mu_0 \left( \mathbf{j} + \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} \right)$$
$$\nabla \times (\nabla \times \mathbf{E}) = -\nabla \times \frac{\partial \mathbf{B}}{\partial t}$$
$$= -\frac{\partial}{\partial t} (\nabla \times \mathbf{B})$$
$$= -\mu_0 \varepsilon_0 \frac{\partial^2 \mathbf{E}}{\partial t^2} - \mu_0 \frac{\partial \mathbf{j}}{\partial t}$$

Maxwell's Equations

$$nonlinear (!)$$

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{u}) = 0$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{j} = 0 \quad \text{``cold''} \quad \text{(plus ``collisions)}$$

$$nm \left[ \frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} \right] = nq (\mathbf{E} + \mathbf{u} \times \mathbf{B})$$

$$\text{``magnetized'' or ``unmagnetized''} \quad \text{``Wasking}$$

$$nm \left( \frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} \right) = nq (\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla$$

Plasma Fluid Dynamics



### **Review of EM Waves in Medium**

$$\varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} + \mathbf{j} = \varepsilon_0 \overline{\varepsilon}(\omega) \frac{\partial \mathbf{E}}{\partial t} \qquad \mathbf{j}(\omega) = \mathbf{j}(\omega) \mathbf$$

$$\begin{cases} \mathbf{k}\mathbf{k} - k^2 \mathbf{I} + \frac{\omega^2}{c^2} \mathbf{I} \\ \\ \mathbf{k}\mathbf{k} - k^2 \end{cases}$$

 $= \bar{\sigma}(\omega) \cdot \mathbf{E}(\omega)$ 



All of the plasma physics here





**EM Waves with**  

$$\varepsilon_{0} \frac{\partial \mathbf{E}}{\partial t} + \mathbf{j} = \varepsilon_{0} \overline{\varepsilon}(\omega) \frac{\partial \mathbf{E}}{\partial t} \qquad \mathbf{j}(\omega) =$$
For High FREQUENCY LADVES,  
COVINENT  

$$\overline{J} = -em_{0}$$

$$m_{e} \frac{dv}{dv} = -eE \Rightarrow$$

$$\overline{J} = \mathbf{j} \frac{e^{2}m}{m_{e}} \frac{1}{m_{e}}$$

$$\overline{J} = \mathbf{j} \varepsilon_{0} (w_{e}^{2})$$

$$\overline{\zeta} = \mathbf{j} \varepsilon_{0} \omega \begin{pmatrix} w_{e}^{2} \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

$$\overline{\varepsilon} = \overline{\mathbf{I}} + \frac{i}{\omega \varepsilon_{0}} \overline{\varepsilon} =$$

out Magnetic Field  $= \overline{\sigma}(\omega) \cdot \mathbf{E}(\omega) \qquad \mathbf{\varepsilon}_{\iota}$   $\omega \neq \omega_{\mathbf{p}_{\iota}}, \text{ THEN ONLY ELECTRON}$  $\boldsymbol{\varepsilon}_{\omega} = \mathbf{I} + \frac{\mathbf{I}}{\omega\varepsilon_0} \boldsymbol{\sigma}_{\omega}$ ve  $-j\omega \overline{\nabla} = -\frac{e}{m_0} \overline{E}$   $\overline{\nabla} = \cdot \left(-\frac{e}{m_0}\right) \overline{E}/\omega$ Ē  $\overline{E}$  $\begin{array}{c} \begin{array}{c} \begin{array}{c} \begin{array}{c} & 0 \end{array} \\ & & \\ \end{array} \end{array} \end{array} \end{array} \end{array} \end{array} \end{array} \left( \begin{array}{c} \\ Fon \end{array} \overline{E} = \left( \begin{array}{c} E_{x}, E_{y}, E_{z} \right) \\ \\ \end{array} \right) \\ \begin{array}{c} \\ \end{array} \end{array}$  $\overline{I}\left(1-\frac{\omega^{2}}{2}\right)^{2}$ 



### "Cold" EM Waves without Magnetic Field FON ELECTROMAGNETIC LAVES TAKE h= Zh on N=ZN $\boldsymbol{\varepsilon}_{\omega} = \mathbf{I} + \frac{1}{\omega\varepsilon_0}\boldsymbol{\sigma}_{\omega}$ $\begin{cases} f - N^2 - \frac{4\omega^2}{\omega^2} & 0 & 0 \\ 0 & 1 - N^2 - \frac{\omega^2}{\omega^2} & 0 \\ 0 & 0 & 1 - \frac{\omega^2}{\omega^2} & 0 \\ 0 & 0 & 1 - \frac{\omega^2}{\omega^2} & 0 \\ \frac{\omega^2}{\omega^2} & \frac{\omega^2}{\omega^2} & \frac{\omega^2}{\omega^2} \\ \frac{\omega^2}{\omega^2} & \frac{\omega^2}{\omega^2} & \frac{\omega$ \*\*\* (!!) FARADAR'S LAN POLANIZATION: hxE=wB ELL TRANSVERSOWAVES ELLA Long, TUDINA MAUTI EULA So B=o For Zong, TUDINAL GALES





### "Warm" EM Waves without Magnetic Field

#### WITH ELECTION PRESSORE ...

WHEN A. V70 => COMPRESSION/DECOMPRESSION

For Longituping WAUES

on

Ans

 $\Rightarrow$ 

ージャン====-ジャ<u>たん</u>  $-j\omega \overline{\upsilon} \left( 1 - \gamma \beta_{2} \beta_{3} \frac{\omega^{2}}{\omega^{2}} \right) = \overline{m} \overline{E}$  $\overline{J} = q m \overline{J} \Longrightarrow \qquad 6 = 3 \omega \epsilon_0 \frac{\omega_p^2 / \omega_v^2}{1 - v_h^2 v_h^2 (\omega_p^2 / \omega_v^2)}$ 

W= W(1+85%) ELECTION Rutsan Cutes

WHEN h.V=O => NO CHANGO OF DENSITY

### Waves in Magnetized Plasma





 $\hat{v}^{\pm} = \hat{v}_x \pm i\hat{v}_y$ ,  $\hat{E}^{\pm} = \hat{E}_x \pm i\hat{E}_v$ 

$$\hat{v}^{\pm} = i \frac{q}{\omega m} (\hat{E}^{\pm} \mp i \hat{v}^{\pm})$$

$$(\hat{E}_y - \hat{v}_x B_0)$$

$$\hat{v}^{\pm} = i\frac{q}{m}\hat{E}^{\pm}\frac{1}{\omega\mp s\omega}$$





# Waves in Magnetized Plasma







(These formula will be on your quiz!)



$$\boldsymbol{\varepsilon}_{\omega} = \mathbf{I} + \frac{\mathbf{i}}{\omega\varepsilon_{0}}$$
$$\boldsymbol{\varepsilon}_{(\omega)} = \begin{pmatrix} S & -\mathbf{i}D & 0 \\ \mathbf{i}D & S & 0 \\ 0 & 0 & P \end{pmatrix}$$

Working both in the laboratory and with theoretical calculations, he found many ways to put waves to work in fusion research in succeeding decades, and his 1962 book, "The Theory of Plasma Waves," codified the subject in mathematical form for the first time.

http://www.nytimes.com/2001/04/18/nyregion/thomas-h-stix-plasma-physicist-dies-at-76.html

Waves in Magnetized Plasma  $\boldsymbol{\sigma}_{\boldsymbol{\omega}} \qquad S = 1 - \sum_{\alpha} \frac{\omega_{p\alpha}^2}{\omega^2 - \omega_{c\alpha}^2}$  $D = \sum_{\alpha} s_{\alpha} \frac{\omega_{p\alpha}^2}{\omega^2 - \omega_{c\alpha}^2} \frac{\omega_{c\alpha}}{\omega}$  $P = 1 - \sum_{\alpha} \frac{\omega_{p\alpha}^2}{\omega^2}.$ 



$$S = 1 - \sum_{\alpha} \frac{\omega_{p\alpha}^2}{\omega^2 - \omega_{c\alpha}^2}$$

$$D = \sum_{\alpha} s_{\alpha} \frac{\omega_{p\alpha}^2}{\omega^2 - \omega_{c\alpha}^2} \frac{\omega_{c\alpha}}{\omega}$$

$$P = 1 - \sum_{\alpha} \frac{\omega_{p\alpha}^2}{\omega^2}.$$

$$R \equiv 1 - \sum_{s} \frac{\omega_{ps}^2}{\omega^2} \left(\frac{\omega}{\omega \pm \Omega_s}\right)$$

$$L \equiv 1 - \sum_{s} \frac{\omega_{ps}^2}{\omega^2} \left(\frac{\omega}{\omega \mp \Omega_s}\right)$$

$$S \equiv \frac{1}{2} (R + L) \quad D \equiv \frac{1}{2} (R - L)$$

$$P \equiv 1 - \sum_{s} \frac{\omega_{ps}^2}{\omega^2}. \qquad \boldsymbol{\varepsilon}(\omega) =$$



### THE THEORY OF PLASMA WAVES



THOMAS HOWARD STIX Professor of Astrophysical Sciences, Princeton University

The Advanced Physics Monograph Series McGraw-Hill Book Company

# $= \begin{pmatrix} S & -iD & 0 \\ iD & S & 0 \\ 0 & 0 & P \end{pmatrix}$



### Clemmow-Mullaly-Allis 3 $\Omega_i = \omega$ 2 $|\Omega_e|/\omega$ Zm.e -Lecture 12: CMA Diagram $|\Omega_e| = \omega$ R = 0"X-mode catal (0) X for 0=== $\mathbf{U} = \boldsymbol{\omega} \mathbf{k} / \mathbf{k}^2 \mathbf{C}$ 0

0



#### **Recent milestones from the APEX Collaboration, on** the path toward confined e+e- pair plasmas

Eve V Stenson (Max Planck Institute for Plasma Physics)



#### ... and significantly increases the number of available modes, even in simple systems.



![](_page_12_Picture_8.jpeg)

 $\mathcal{N} = \frac{kc}{\omega}$ 

 $\begin{cases} S - \mathcal{N}^2 \cos^2 \psi & -iD \\ iD & S - \mathcal{N}^2 \\ \mathcal{N}^2 \cos \psi \sin \psi & 0 \end{cases}$ 

 $\mathbf{k} = (k \sin \psi, 0, k \cos \psi)$ 

Good places to start: Propagation along B ( $\psi = 0$ ) Propagation  $\perp$  to B ( $\psi = \pi/2$ )

### Waves in Magnetized Plasma

$${\begin{array}{*{20}c} \mathcal{N}^{2}\cos\psi\sin\psi\\ {}^{2}&0\\ P-\mathcal{N}^{2}\sin^{2}\psi \end{array} } \cdot \begin{pmatrix} \hat{E}_{x}\\ \hat{E}_{y}\\ \hat{E}_{z} \end{pmatrix} = 0$$

![](_page_13_Figure_7.jpeg)

![](_page_14_Figure_0.jpeg)

![](_page_15_Figure_2.jpeg)

om NASA GSFC, adapted from Thorne et al. 2005

Tsurutani and Smith, 1974

![](_page_16_Picture_0.jpeg)

### Extra-Ordinary Mode

$$\begin{pmatrix} S & -iD \\ iD & (S - \mathcal{N}^2) \end{pmatrix} \cdot \begin{pmatrix} \hat{E}_x \\ \hat{E}_y \end{pmatrix} = 0$$
$$\mathcal{N}_X = \left(\frac{S^2 - D^2}{S}\right)^{1/2}$$

$$\omega_{\rm uh} = (\omega_{\rm ce}^2 + \omega_{\rm pe}^2)^{1/2}$$

S = 0 at the "upper hybrid resonance" and the "lower hybrid resonance" w ions

### Waves k L B

### Plus: "Ordinary" Mode

$$\omega^2 = \omega_{\rm pe}^2 + k^2 c^2$$

Ordinary Mode only depends upon electron density  $\Rightarrow$ Interferometry (!)

![](_page_16_Picture_10.jpeg)

![](_page_17_Picture_0.jpeg)

### Extra-Ordinary Mode

$$\begin{pmatrix} S & -iD \\ iD & (S - N^2) \end{pmatrix} \cdot \begin{pmatrix} \hat{E}_x \\ \hat{E}_y \end{pmatrix} = 0$$

![](_page_17_Figure_3.jpeg)

### Waves k L B

### Plus: Ordinary Mode

$$\omega^2 = \omega_{\rm pe}^2 + k^2 c^2$$

# **Example Problems and Discussion**

- Plasma waves with electron flow
- Finding simple limits to plasma wave dispersion
- Plasma wave damping due to collisions ( $\omega = \omega_r + i \omega_i$  or  $k = k_r + i k_i$ )
- Whistler waves
- Interferometry and Faraday Rotation

### Part 6: Two beams and a Plasma (25 points total)

Consider a finite-sized plasma into which two counter-streaming beams are injected.

![](_page_19_Figure_2.jpeg)

Each beam has the same velocity,  $V_0$ , and the same plasma density, and the density of the stationary plasma is equal to the densities of each beam.

The dispersion relation for electrostatic plasma waves in this system is given by

 $D(\omega,k) = 1 -$ 

constant of the order of 3.

$$-\frac{\omega_p^2}{\omega^2} - \frac{\omega_p^2}{\left(\omega - kV_0\right)^2} - \frac{\omega_p^2}{\left(\omega + kV_0\right)^2}$$

If we assume that the longest wavelength that can exist within the plasma is given by the plasma length, L, then prove that two-stream instabilities are stabilized when the density is sufficiently small. Indeed, the stability requirement is  $\omega_p/V_0 < K/L$ , where K is a

Consider a uniformly magnetized and a fully-ionized plasma made from carbon. There would be six times the density of electrons than of ions (but the plasma would still be approximately charge-neutral.)

Describe the Alfvén wave, the electron whistler wave, and the ordinary wave in this plasma. How do these waves (in a fully-ionized carbon plasma) compare to the same waves in a plasma made from singly-ionized carbon having the same mass density (*i.e.* the same density of carbon)?

### Neglecting plasma resistivity, Ohm's Law is flow velocity.

Using Maxwell's equations, derive the dispersion relation for a shear Alfvén wave, propagating along an otherwise unperturbed, straight magnetic field,  $\mathbf{B}_0$ . Consider only perturbations such that the perturbed plasma motion,  $\mathbf{V}_1$ , and the perturbed magnetic field,  $\mathbf{B}_1$ , both perpendicular to  $\mathbf{B}_0$ .

Describe the polarization of the Alfvén wave by relating the perturbed electric field and plasma velocity to the perturbed magnetic field. Are these perturbations oscillating in-phase, out-of-phase, or with a phase-shift?

Neglecting plasma resistivity, Ohm's Law is  $\mathbf{E} + \mathbf{V} \times \mathbf{B} = 0$ , where V is the plasma mass

### Show that in the limit $\omega \to 0$ ,

![](_page_22_Figure_1.jpeg)

Consider the effect of electron collisions on plasma waves in a uniform cold plasma. The equation for electron motion is

$$m_e \frac{d\mathbf{v}_e}{dt} = -e\mathbf{E}$$

where **E** and **B** are the electric and magnetic field of the electromagnetic wave. (There is no equilibrium magnetic field in this problem.)

#### Part A

Show that the effect of collisions can be represented by the substitution

 $m_e \rightarrow m$ 

where  $\omega$  is the wave frequency.

#### Part B

Find the linear dispersion relation for longitudinal electron plasma oscillations including the effects of collisions. Briefly discuss the dissipation of these oscillations when  $\nu \ll \omega_{pe}$ .

#### Part C

imaginary parts of the wave number  $(k = k_r + ik_i)$  when  $\nu \ll \omega$  and  $\omega_{pe} \ll \omega$ .

 $-e\mathbf{v}_e \times \mathbf{B} - m_e \nu \mathbf{v}_e$ 

$$m_e\left(1+\frac{i\nu}{\omega}\right)$$

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For electromagnetic waves in a cold plasma, find approximate expressions for the real and
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The whistler wave propagates along the magnetic field with a frequency range,  $\omega_{ci} \ll \omega < \omega$  $\omega_{ce}$ . An approximate dispersion relation for the whistler wave is

 $\frac{k^2c^2}{\omega^2} =$ 

where the sign of all frequencies are positive in the expression above. Consider the propagation of whistler's in the magnetospheres of planets caused by atmospheric lightning. At the moment of the lightning strike, a broad band of electromagnetic frequencies are excited in a single pulse. Some of this wave energy couples to whistler waves that propagate from one pole to the other. Because the waves have different group velocities, waves at different frequencies arrive at different times.

#### Part A

Describe the steps needed to derive the whistler wave dispersion relation in a cold magnetized plasma.

### Part B

whistler pulse as a function of the wave frequency.

### Part C

At a "cut-off" frequency, the wavelength becomes very long. At a "resonance", the wavelength tends toward zero. Is there a "cut-off" or "resonance" for whistler waves? If so, what happens at this "cut-off" or "resonance"?

$$=\frac{\omega_{pe}^2}{\omega(\omega_{ce}-\omega)}$$

Since the group velocity is  $v_q = \partial \omega / \partial k$ , given an expression for the time of arrival of

![](_page_25_Figure_0.jpeg)

Fig. 6.5 (a) Laser interferometer in Mach-Zehnder arrangement, (b) microwave interferometer. The optical arrangement uses partially-reflecting and fully-reflecting mirrors. The analog to a par-tially reflecting mirror is the directional coupler for microwaves the quantitative relationship between  $\Delta\phi$ and the plasma?

**Partice 6.1** Cut-off densities for microwave and laser interferometers

plasma condition that average beration of the use of the second s Source λ

### Mcartwave

3 cm

HCN-laser Under what conditions will  $\Delta \phi$  change? 890 GHz CO<sub>2</sub> laser 10.6 µm

![](_page_25_Figure_7.jpeg)

The interferometer only works provided the plasma density is not too high. Describe the Cut-off-density  $n_{\rm co}({\rm m}^{-3})$  $1.2 \times 10^{18}$ 10 GHz If a magnetic field is **applie**d perpendicular **37 GHz** ath of the optila beal Owhile keeping all other average plasma parameters constant GHZ what conditions will  $9\Delta\phi$  remain 1071

7 890'GHz	$9.8 \times 10^{21}$
268 THz	$9.9 \times 10^{24}$

In this problem, you are to derive expressions for the phase shift of the laser passing through the plasma as described in Question 6.

For typical magnetized laboratory plasmas, the laser frequency is much greater than the electron plasma and cyclotron frequencies. With the z-axis aligned with the magnetic field, the plasma dielectric tensor is

$$\overline{\overline{\epsilon}} = \left(\begin{array}{ccc} S & -iD & 0\\ iD & S & 0\\ 0 & 0 & P \end{array}\right)$$

where

$$S = 1 - \frac{\omega_{pe}^2}{\omega^2 - \omega_{ce}^2}; \quad D = -\left(\frac{\omega_{ce}}{\omega}\right) \frac{\omega_{pe}^2}{\omega^2 - \omega_{ce}^2}; \quad \text{and } P = 1 - \frac{\omega_{pe}^2}{\omega^2}.$$

Maxwell's equations describe the electromagnetic wave of the laser with the three equations

$$\left[\mathbf{nn} - n^2 \overline{\overline{I}} + \overline{\overline{\epsilon}}\right] \cdot \mathbf{E} = 0 \,,$$

where  $\mathbf{n} = \mathbf{k}c/\omega$  is the vector index of refraction.

### Interferometry

• For the first case, the laser beam propagates *across* the plasma cylinder and *perpendicular* to the magnetic field. For this case, the total change in the phase-shift of the laser beam integrated across the plasma is

$$\Delta \phi = \int dx \, (n-1)(\omega/c) \, .$$

Find the correct expression for the phase shift when  $\omega \gg \omega_{pe}$ .

• For the second case, the laser beam propagates *along* the plasma cylinder and *parallel* to the magnetic field. For this case, an initially linearly polarized laser beam must be decomposed into the two propagating electromagnetic waves that propagate along **B**. These waves are called the "right-hand circularly polarized" and "right-hand circularly polarized" waves.

Find expressions for the polarizations of the right-hand and left-hand waves and show that the indices of refraction for these waves are (right-hand)  $n^2 = S + D =$  $\frac{1 - \omega_{pe}^2}{\omega(\omega - \omega_{ce})} \approx 1 - (\omega_{pe}^2/\omega^2)(1 + \omega_{ce}/\omega + \dots) \text{ and (left-hand)} \quad n^2 = S - D = 1 - \omega_{pe}^2/\omega(\omega + \omega_{ce}) \approx 1 - (\omega_{pe}^2/\omega^2)(1 - \omega_{ce}/\omega - \dots).$ 

What is the relative difference per unit distance of propagation between the phase shifts of the right and left-handed waves?

How quickly does the polarization of the laser beam change per unit distance of propagation when  $\omega \gg \omega_{pe}$  and  $\omega \gg \omega_{ce}$ .

### Faraday Rotation

![](_page_26_Figure_20.jpeg)

![](_page_26_Figure_21.jpeg)

![](_page_26_Figure_22.jpeg)

![](_page_26_Figure_23.jpeg)

![](_page_26_Figure_24.jpeg)

![](_page_26_Figure_25.jpeg)

Consider a magnetized plasma cylinder of radius a. You are to measure the line-density of the plasma by measuring the phase change of a polarized laser beam passing through the plasma.

- the density along the laser's path?
- phase-shift is proportional to the integral of the density along the laser's path?
- along the axial magnetic field in the  $\hat{z}$ -direction?

• If the plasma is confined by an axial plasma current so that the magnetic field is entirely azimuthal (*i.e.*  $\mathbf{B} = \theta B_{\theta}$ , of the so-called z-pinch), how must the laser be polarized so that the interferometer's phase-shift is proportional to the integral of

• If the plasma is confined by a magnetic field that is directed along the  $\hat{z}$ -direction (*i.e.* the so-called  $\theta$ -pinch), how must the laser be polarized so that the interferometer's

• For the  $\theta$ -pinch (with the magnetic field directed along the  $\hat{z}$ -direction), what happens to the polarization of a laser beam, which is linearly polarized and propagating