Lecture 19: Beam-Plasma Instability Plasma Physics 1

APPH E6101x **Columbia University**

Instabilities

- "Micro-instabilities"
 - Velocity-space
 - Drift wave instabilities
- "Macro-instabilities"

• ...

- Gravitational
- Kink, Sausage, etc
- Ballooing

•

• Resistive, tearing modes



This Lecture (Non-Relativistic, Cold-Plasma) Electron-Beam Plasma Instability

- Linearized electron fluid equations
- **Dispersion relation**
- Complex eigenfrequencies, wavenumber
- Energetics









A. Lvovskiy et al 2020 Nucl. Fusion 60 056008

Observations of the Beam-Plasma Instability

K. W. GENTLE AND C. W. ROBERSON* The University of Texas at Austin, Austin, Texas 78712 (Received 30 June 1971)

The nonlinear limit of the instability driven by a low density cold electron beam in a collisionless plasma is experimentally found to be determined by the trapping of the beam by the most rapidly growing wave.

The apparatus has previously been described,⁵ where ficient to render the dynamics one dimensional and the it was used for a similar experiment to test quasilinear background pressure of 8×10^{-6} Torr precluded collisional effects. The plasma is quiet, with low-frequency theory. The apparatus and techniques are also very density fluctuations of only a few percent. The axial similar to those used by Malmberg and Wharton⁶ for density uniformity is excellent. Although the density the problem. The present work complements and exdrops approximately 20% over the 40 cm near the gun, tends the results of that paper in the nonlinear regime. no measurable gradient exists over the remainder of the For these experiments, the two-meter column contained column. A gradient of more than a few percent would a 3-cm diam plasma with a density near 10° and a tembe readily detectable. The density gradient near the gun perature of 20 eV. The magnetic field of 1 kG was sufis not significant. The waves are still far from saturation when they enter the uniform region, and all the important physics occurs in that region.

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Typical results for the development of the instability are shown in Fig. 1. The wave power is measured with a loosely coupled, calibrated (-32 dB) coaxial probe connected to a broad band amplifier and rf voltmeter. The probe may be moved the length of the machine, and the figure shows the result for the absolute power in the waves. The qualitative behavior is precisely that predicted: linear growth to a peak, followed by slow oscillation. Although the theory was presented for infinite geometry and an initial value problem, the argument can easily be applied to finite geometry and growth in space. The peak wave power should be

$$P_w = 2^{2/3} \eta^{1/3} P_B, \tag{1}$$

where $P_b = I_b V_b$, the input beam power, and

$$\eta = \int_{0}^{a} n_{b}(r) r \, dr \left(\int_{0}^{a} n_{p}(r) r \, dr \right)^{-1}. \tag{2}$$



FIG. 1. Total wave power as a function of distance from the point of injection of the beam into the plasma column. The background plasma had a density of 7×10^8 and a temperature of 18 eV.



FIG. 2. Total wave power measured at the spatial maximum as a function of beam current. The wave power is in relative units.

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tions in the potential well. In space, the oscillations will appear with a wavelength determined by the particle oscillation frequency in the well and the propagation velocity, the beam velocity u. If the potential is sinusoidal and the particles are trapped near the bottom of the well, the oscillation length is given by

$$\lambda_{\rm osc} = 2\pi u (m/ekE)^{1/2}. \tag{3}$$

Making use of Eq. (1), we note that this implies that



$$\lambda_{\rm osc} \propto I_b^{-1/3}. \tag{4}$$

FIG. 3. Oscillation wavelength of the wave energy as a function of beam current.

NONLINEAR EVOLUTION OF A TWO-STREAM INSTABILITY*

K. V. Roberts[†] and H. L. Berk University of California, San Diego, La Jolla, California (Received 12 June 1967)

Calculations of a two-stream instability have been made by following the motion of the phase-space boundaries of an incompressible and constant-density phase-space fluid. Because of the condensation of holes, which to a good approximation act as gravitational particles, large-scale nonlinear pulses develop.

$$\frac{df}{\partial t} + v \frac{\partial f}{\partial x} - \frac{\partial \varphi}{\partial x} \frac{\partial f}{\partial v} = 0.$$

$$\frac{dx}{dt} = v, \quad \frac{dv}{dt} = -\frac{\partial \varphi}{\partial x},$$

$$\frac{\partial^2 \varphi}{\partial x^2} = \omega_p^2 \left[\int f \frac{dv}{v_0} - 1 \right].$$

The example to be discussed is a two-stream are $v_0 \Delta t / \Delta x = 0.25$, $\omega_D \Delta t = 1/20$, and $\Delta x = L/64$, instability, in which the electron plasma is where Δx is the grid used for evaluating Poisslightly perturbed at t = 0 from an equilibrium son's equation. The unstable wave numbers characterized by four straight lines in phase are $k = 2\pi n/L$ with n = (1, 2), and the linear space: f = 1 for $\frac{1}{2}v_0 < |v| < v_0$ and f = 0 elsewhere. growth rates are $\gamma/\omega_{b} = 0.30, 0.315$. Periodic boundary conditions are imposed at x = (0, L) and the parameters of the problem

The most striking feature of the calculation is the behavior of the f = 0 "cavity" which initially occupies the strip $(|v| < \frac{1}{2}v_0)$ between the two plasma layers. This must preserve constant area as it deforms, and it is seen in Fig. 1 to coalesce into holes of roughly elliptical shape, so that a large-amplitude electrostatic wave is set up. Superimposed on this wave are coherent oscillations due to rotation of the holes in phase space, and also random fluctuations due to the motion of smaller elements of the hole "fluid." The two outer curves adjust almost adiabatically to the instantaneous potential function.



FIG. 1. Evolution in phase space of a two-stream instability from time steps 200 to 600 at intervals of 50 steps. Each step is 1/20 of a plasma period, and the horizontal and vertical coordinates are x, v, respectively. Periodic boundaries have been imposed and three identical periods are shown along each row. The shaded area represents the f=0 region enclosed by the plasma fluid.



Linearize Beam-Plasma

ELECTRON DEAN, MO VO CLINEARIZE FOR ELECTHOSTATIC CULD PLASMA, MP JUNEARIZE FOR ELECTHOSTATIC STATIONARY long M. = M+Mp • CONTINUITY: $\frac{\partial m}{\partial L} +$ $MOMENTUM: -j \omega V = -\frac{9}{m}$ -jw~p+jk.V, • POISSON'S EQUATION: $-\frac{1}{2}\vec{\Phi} = \tilde{n} = \frac{m}{r_0} \frac{\bar{k} \cdot \bar{V}_p}{G} = \frac{\bar{V}_p}{F} = -\frac{3}{m_0} \frac{\bar{E}}{G} = -\frac{3}{m_0} \frac{\bar{E}}{G}$

Linear Beam-Plasma Dispersion Relation

LINEAN DISPERSION PELATION

 $D(s, w) \overline{\Phi} = 0$

00

A BEAM-MASMA SESTEM HAS THREE (3) 1200TS Ard Joy nhust un Waho - Wn Wiw

 $-k\hat{Q} = -\frac{\pi^2}{\epsilon_0} \left[\frac{k\hat{Q}}{p_0} + \frac{k$

 $D(S, \omega) = (-\frac{\omega^2}{6^2} - \frac{\omega^2}{6^2})^2$

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Energetics of Linear Modes

EXAMINE THESE MODES

FLEST STEP: WHAT ANG ME WALF ENENgy FOR

THE "FACT BEAM" LIMIT RUDGE WING

WHAT IS THE TIME-AVENAGES "WAVE ENERgy"?

 $\left(\text{LARE ENERGY} \right) = \frac{1}{2} \epsilon_0 \left| \vec{E} \right|^2 + \frac{1}{2} \left(\frac{m_0 + \vec{m}}{m_0} \right) m_0 \left(\frac{v_0 + \vec{v}}{m_0} \right)^2 \right| \epsilon_{imor}$

 $z = f_0 \left[\overline{E} \right] + \frac{1}{2} M_{p} \left[m_0 \right] p$

 $t \pm m_{bo} m_{b} \left[\tilde{V}_{b}^{2} \right] + m_{b} V_{b} \left[\tilde{m} \tilde{V}_{b} \right] t - -$

Energetics of Plasma Wave

WARE #1 : PLASMA WAVE WEWP $\tilde{V}_{p} = -j \frac{\tilde{E}}{n_{0}} \frac{\tilde{E}}{\omega_{p}} \qquad \tilde{m}_{p} = -j \frac{\tilde{E}n_{p}}{n_{0}} \frac{\tilde{h}_{i}\tilde{E}}{\omega_{p}}$ $V_0 = -j \frac{E}{m_0} \left(-hu\right)$ $\tilde{M} = -j \frac{m_0}{m_0} \left(-hu\right)^2$ (m)2 201 PLASMA WACRS *** HAVE TWICO (24) ENENGY OF ELETTIC FIELD

Energetics of "Fast" Beam Mode

WAVE #Z: FAST JEAN NODE WITH +240

V = - j Moho mp = - j moho h.E Bitt Ano SMALL D'Enteurs Works

 $\overline{V} = -j \frac{g}{m_0} \frac{E}{\omega}$ $\overline{N} = -j \frac{g}{m_0} \frac{g}{\omega} \frac{g$

 $\left(\text{WARE ENERGY} \right) = \frac{c_0}{2} \left| \overline{E} \right|^2 + \frac{c_0}{2} \left| \overline{E} \right|^$

SO FAST BEAM MODO >> 2×(E)E/2)

Energetics of "Slow" Beam Mode

WAVE #3 "SLOW BEAN NOOD WERE AUD

Up - SmAcc $\tilde{V} = -\frac{\pi}{m_0} \frac{\tilde{E}}{[-\omega_0]} \qquad \tilde{m}_0 = -j \frac{gm_0}{m_0} \frac{k_c \tilde{E}}{[\omega_0^2]}$ 5001

 $\left(\left(u \, A \, u \, \varepsilon \, E \, n \, \varepsilon \, n \, g \, \tau \right) = \frac{\varepsilon_0}{2} \left| \tilde{\varepsilon} \right|^2 + \frac{\varepsilon_0}{2} \frac{\omega_0}{\omega_1} \left| \tilde{\varepsilon} \right| - \frac{\omega_0}{2} \frac{|\tilde{\varepsilon}|^2}{\omega_1} \left| \tilde{\varepsilon} \right| - \frac{1}{2} \frac{|\tilde{\varepsilon}|^2}{\omega_1} \left| \tilde{\varepsilon} \right|^2$

WITH hos >>1, THE "WARE ENERgy" IS 3 STIR ENERGY IS ***

(WHAT DOES NEGATIVE ENERGY MEAN?)

Dispersion Diagram of E-Beam Plasma System



INSTABILITY DECORS WHEN

POSITIVE ENENgy PLASMO Moos

NEGATUS ENENGY *** BEAN ~ MODE

Dispersion Diagram of E-Beam Plasma System

RECONATION JETHEEN PLASMA + BEAM:

DEFINE WE Up + DW









8.1.5 Temporal or Spatial Growth

$$kv_0 = \omega + i \left(\frac{\alpha_b \omega_{pe}^2 \omega^2}{\omega_{pe}^2 - \omega^2}\right)^{1/2}$$

$$k_{\rm I} = \frac{\omega_{\rm pe}}{v_0} \frac{\alpha_{\rm b}^{1/2} \omega}{(\omega_{\rm pe}^2 - \omega^2)^{1/2}}$$

These waves are convectively unstable: growth as wave propagates at V₀



Summary

- E-Beam Plasma is unstable when $kV_0 \sim \omega_p$
- "Slow" beam mode is a "negative energy" wave
- Electric field energy increases as unstable wave grows
- Instabilities result from a source of energy and when a wave, or displacement, releases the source of energy
- NEXT LECTURE: ideal "macroinstability", flute-interchange mode, and kink mode

ANNALS OF PHYSICS: 1, 120–140 (1957) Stability of Plasmas Confined by Magnetic Fields¹

M. N. Rosenbluth* and C. L. Longmire

Los Alamos Scientific Laboratory, University of California, Los Alamos, New Mexico

In this paper, we examine the question of the stability of plasmas confined by magnetic fields. Whereas previous studies of this problem have started from the magnetohydrodynamic equations, we pay closer attention to the motions of individual particles. Our results are similar to, but more general than, those which follow from the magnetohydrodynamic equations.

that the internal energy of the plasma per unit mass is

 $E_p =$

where p is the pressure and v the specific volume. In any adiabatic motion,

$$=\frac{pv}{\gamma - 1}$$
(22)

$$\sim v^{-\gamma}$$
 (23)

Laboratory Magnetospheres: Designed for Maximum Flux Tube Expansion

Levitated Dipole Experiment (LDX) **Flux Tube Expansion:** $\delta V(out)/\delta V(in) = 100$

3.6 m



$$V = \int \frac{dl}{B} \propto L^4$$



Ring Trap 1 (RT-1) **Flux Tube Expansion:** $\delta V(out)/\delta V(in) = 40$