Lecture 18: Plasma Physics 1

APPH E6101x Columbia University

Last lecture: Reduced MHD and Kink Modes This lecture: Numerics, Perturbed Energy, Tearing Modes

Ch. 9 Fitzpatrick: Magnetic Reconnection

Magnetic Reynolds Number Lundquist Number

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{V} \times \mathbf{B}) + \frac{\eta}{\mu_0} \nabla^2 \mathbf{B}.$$

$$\begin{split} \frac{\partial \mathbf{D}}{\partial t} &= \nabla \times (\mathbf{V} \times \mathbf{B}) + \frac{\eta}{\mu_0} \nabla^2 \mathbf{B}. & \nabla \cdot \mathbf{V} = 0, \quad \mathbf{V} = \nabla \phi \times \mathbf{e}_z, \\ \rho \left[\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right] + \nabla p - \mathbf{j} \times \mathbf{B} = 0, \quad \mathbf{B} = \nabla \psi \times \mathbf{e}_z, \\ S &= \frac{\mu_0 V L}{\eta} \simeq \frac{|\nabla \times (\mathbf{V} \times \mathbf{B})|}{|(\eta/\mu_0) \nabla^2 \mathbf{B}|}, \quad \rightarrow \infty \text{ ("Ideal")} \quad \mathbf{E} + \mathbf{V} \times \mathbf{B} = \eta \mathbf{j}. \quad \approx 0 \text{ ("Ideal")} \end{split}$$

$$S = \frac{\tau_R}{\tau_H}$$
. $\tau_H = \frac{k^{-1}}{(B_0^2/\mu_0 \rho)^{1/2}}$ $\tau_R =$

Reduced MHD: Incompressible

$$\mu_0 a^2$$



 $\mu_0 J_z = \nabla^2_\perp \psi$

Summary of Reduced MHD

 $\overline{V_{1}}(2, \sigma_{1}^{2} + 1 = 2 \times \nabla X$ $\overline{B}_{f}(n, \sigma_{r}^{2} \epsilon) = \overline{\mathcal{F}} \times \nabla \mathcal{V}$ 0.0=0 $\nabla \cdot \overline{B} = 0$

POLOO FLUX EVOLUES DY~AMICALY ٦Ł DUB TO FIELD-ALIENED CHANGES IN THE STREAM FUNCTION "



13 AXIAC VORTICITY CFIANGES According 70 FIELD-ALIGNES VARCATION OF AXIM COPRENT "



Linearized Reduced ("Ideal") MHD



$$\frac{J_{z,0}}{\partial r}\tilde{\psi} + \frac{B_p}{\mu_0 r}\left(m - nq\right)\nabla_{\perp}^2\tilde{\psi}$$

Shafranov "Constant J" Case $(dJ_0/dr = 0)$ Wesson "Model $J_0(r)$ " Case $(dJ_0/dr \neq 0)$

Wesson's Cylindrical Equilibrium $(dJ_0/dr \neq 0)$

J.A. Wesson 1978 Nucl. Fusion 18 87; http://doi.org/10.1088/0029-5515/18/1/010

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John Weston 1932 - 2020





Jo = CENMA CURRENT DENSITY $g(u) = \frac{2B_{2}}{M_{0}RJ_{0}} \qquad M_{0}T_{p} = \frac{2\pi^{2}q^{2}B_{2}}{M_{0}g(u)R(1+V_{*})}$ $V = \frac{g(u)}{g(u)} - 1$ $(P) \sim \frac{e^{2}}{g^{2}} \qquad B_{p} - 1 \qquad CB_{n} > -\frac{20E}{2}$

g(0), g(a) Two PANAMETERS dg = S = MAGACTIC SHEAN #* F PRESSUR Profils Com DE mora PRAKES



LINEARIZED EQUATIONS FOR PERTURSED STREAM FUNCTION (X) AND PETTURSED PULSION FLUX (4)

LET'S TARK Q= UNIFORM, WITH A SHARP JUMP AT THE PLASMAS EDSE: SUNFACE CURNENT AT RASand's ESgl

WITH INDUCTION EQUATION: ποω Το FIGUNO O JT $\hat{\Psi}(n)$? ; DUT, HOW TO



 $-9\omega\nabla_{1}^{2}\tilde{\chi}=-\frac{m}{n}\frac{2J_{2}}{\partial n}\tilde{\psi}+\frac{B\rho}{\mu_{n}}(m-n\rho)\nabla_{1}^{2}\tilde{\psi}$ pr +10) (induction) $-\omega \hat{\psi} = \frac{R}{2}(m-mq)\tilde{\chi}$



Wesson's Kink Modes









WITH J2(n), WE HAVE TO SOLVE FOR Q USING a COMPUTER. (THIS IS VERY EASY FOR THE CELLIDAICA "TOEdwate",



SINCO [W] < WA, THE KINK MODE CAUSES THE INTERNAL " PLASMA TO RESPOND "QUICKLY", SO QUICKLY THAT WE CAN 190000 THER TIMO CT TARIAS TO FORM A DISTORTED, 3D, QUASI-EQUILIANUM

INSIDE, THE PLASMA IS A "FORCES" EDULIBICA

 $O \approx -\frac{m}{n} \frac{\partial J_z}{\partial n} \mathcal{F} + \frac{Bp}{mon} (m - mg) \mathcal{V}_y^2 \mathcal{F}$

OUTSIDE, THE RESPONSE IS THE "UACUUM" RESPONSE.

F THE SURFACE CURRENT PUSHES PULLS" PLASMA, AND TYTE DISTURTES. PLASMA IS MEASURED BY P(1, 0, 2)



Kink Mode Plasma.nb

OUTSIDE, THE RESPONSE IS THE "UACUUM" RESPONSE.

Eqs for MHD Force Balance

With x = r/a and the Wesson current profile, the equation for MHD force balance is

 $\ln[18] = \exp[deal] = \{D[\psi[x], x] = \psi p[x],$ $x D[x \psi p[x], x] - m^2 \psi [x] +$ $4 v j (m / n / q 0) x^2 (1 - x^)$

$$Dut[18] = \left\{ \psi'[x] = \psi p[x], -m^2 \psi[x] + \frac{4 m x^2 (1 - x^2)^{-1 + vj} v j \psi[x]}{n q_0 \left(-1 + \frac{m (1 - (1 - x^2)^{1 + vj})}{n q_a x^2} \right)} + x (\psi p[x] + x \psi p'[x]) = 0 \right\}$$

INSIDE, THE PLASMA IS A "FORCES" EDUIJBACUN





Kink Mode Plasma.nb



Kink Mode Plasma.nb



Wall "out": b/a = 1.8

Wall "in": b/a = 1.2

$$\rho_{0} \frac{\partial \mathbf{v}_{1}}{\partial t} = \mathbf{j}_{0} \times \mathbf{B}_{1} + \mathbf{j}_{1} \times \mathbf{B}_{0} - \nabla p_{1} = \mathbf{F}(\boldsymbol{\xi})$$

$$\rho_{0} \frac{\partial^{2} \boldsymbol{\xi}}{\partial t^{2}} = \mathbf{F}(\boldsymbol{\xi})$$

$$\underbrace{\frac{\partial^{2} \boldsymbol{\xi}}{\partial t^{2}}}_{2} \int \rho_{0} |\boldsymbol{\xi}|^{2} dV = -\frac{1}{2} \int \boldsymbol{\xi}^{*} \cdot \mathbf{F}(\boldsymbol{\xi}) dV = \delta W(\boldsymbol{\xi}^{*}, \boldsymbol{\xi})$$

$$\omega^2(\boldsymbol{\xi}, \boldsymbol{\xi}^*) = \frac{\delta W(\boldsymbol{\xi}, \boldsymbol{\xi}^*)}{K(\boldsymbol{\xi}, \boldsymbol{\xi}^*)}$$

Sec. 7.3: Introduction to plasma physics by Donald Gurnett and Amitava Bhattacharjee (2nd Ed. 2017; Cambridge) Sec 3.1: Magnetohydrodynamic Stability of Tokamaks by Hartmut Zohm (Wiley 2014)

Energetics of Kink Modes

grated Energy

Perturbed "Potential" Energy Perturbed Kinetic Energy





$$\begin{aligned} & \underset{\rho_{0} \rightarrow v_{1}}{\underset{\partial t}{\partial t}} = \mathbf{j}_{0} \times \mathbf{B}_{1} \\ & \underset{\overline{v}_{2}}{\overset{\omega^{2}}{2}} \int \rho_{0} |\xi|^{2} dV = -\frac{1}{2} \int \xi^{*} \\ & \underset{\varepsilon}{\overset{\omega^{2}}{2}} \int \rho_{0} |\xi|^{2} dV = -\frac{1}{2} \int \xi^{*} \\ & \underset{\varepsilon}{\overset{\varepsilon}{\varepsilon}^{*}} \left(\overline{\mathcal{J}}_{o} \times \overline{\mathcal{I}}_{i} \right) \\ & \underset{\varepsilon}{\overset{\varepsilon}{\varepsilon}} \left(\overline{\mathcal{I}}_{o} \cdot \overline{\mathcal{V}}_{i} \right) - \overline{\nabla} f_{i} \overline{\mathcal{J}}_{o} \\ & \underset{\varepsilon}{\overset{\varepsilon}{\varepsilon}} \left(\overline{\nabla} \chi^{*} \sqrt{\mathcal{V}}_{i} \right) \\ & \underset{\varepsilon}{\overset{\varepsilon}{\varepsilon}} \left(\overline{\nabla} \chi^{*} \sqrt{\mathcal{V}}_{i} \right) \end{aligned}$$

Sec. 7.3: Introduction to plasma physics by Donald Gurnett and Amitava Bhattacharjee (2nd Ed. 2017; Cambridge) Sec 3.1: Magnetohydrodynamic Stability of Tokamaks by Hartmut Zohm (Wiley 2014)

of Kink Modes $+\mathbf{j}_1 \times \mathbf{B}_0 - \nabla p_1 = \mathbf{F}(\boldsymbol{\xi})$

 $\cdot \mathbf{F}(\boldsymbol{\xi}) \, dV = \delta W(\boldsymbol{\xi}^*, \boldsymbol{\xi})$





 $\delta W_p = \frac{1}{2} \int_{r < a} dV \left| \frac{|\mathbf{B}|^2}{\mu_0} + \frac{iJ_{z,0}}{\omega^*} \, \tilde{\mathbf{v}}^* \cdot (\hat{\mathbf{z}} \times \tilde{\mathbf{B}}) \right|$ $= \frac{1}{2} \int_{r < a} dV \left[\frac{|\nabla_{\perp} \tilde{\psi}|^2}{\mu_0} - \frac{iJ_{z,0}}{\omega^*} \hat{\mathbf{z}} \cdot (\nabla_{\perp} \chi^* \times \nabla_{\perp} \tilde{\psi}) \right]$ $\delta W_{v,b} = \frac{1}{2} \int_{a < r < b} dV \frac{|\tilde{\mathbf{B}}|^2}{\mu_0} = \frac{1}{2} \int_{a < r < b} dV \frac{|\nabla_{\perp} \tilde{\psi}|^2}{\mu_0}$ $\delta K = \frac{1}{2} \int_{\mathbb{R} \leq \mathbb{R}} dV \rho \, \frac{|\tilde{\mathbf{v}}|^2}{|\omega|^2} = \frac{1}{2} \int_{\mathbb{R} \leq \mathbb{R}} dV \rho \, \frac{|\nabla_{\perp} \chi|^2}{|\omega|^2} \, .$

$$\omega^2(\boldsymbol{\xi}, \boldsymbol{\xi}^*) = \frac{\delta W(\boldsymbol{\xi}, \boldsymbol{\xi}^*)}{K(\boldsymbol{\xi}, \boldsymbol{\xi}^*)}$$

Energetics of Kink Modes

$$\begin{split} \delta W_p &= \pi m \left[\frac{|\tilde{\psi}_a|^2}{\mu_0} + \frac{J_{z,0}}{\omega^*} (\chi_a^* \tilde{\psi}_a) \right] = \pi m \frac{|\tilde{\psi}_a|^2}{\mu_0} \left[1 - \frac{1}{m} \right] \\ \delta W_{v,b} &= \pi m \frac{|\tilde{\psi}_a|^2}{\mu_0} \Lambda \left(\text{with } \delta W_{v,\infty} = \pi m \frac{|\tilde{\psi}_a|^2}{\mu_0} \right) \\ \delta K &= \pi m \frac{|\tilde{\psi}_a|^2}{\mu_0} \frac{1}{\omega_A^2 (m - nq_a)^2} \,, \end{split}$$

(What happens as $q_a \rightarrow m/n$?)

Perturbed "Potential" Energy Perturbed Kinetic Energy





$$\omega^2(\boldsymbol{\xi}, \boldsymbol{\xi}^*) = \frac{\delta W(\boldsymbol{\xi}, \boldsymbol{\xi}^*)}{K(\boldsymbol{\xi}, \boldsymbol{\xi}^*)}$$

Perturbed "Potential" Energy Perturbed Kinetic Energy

Resistive MHD



"IDEAL" MHO DESCRIBES THE "FATT" TIME DYNAMICS OF A MAGNETIZED PLASMA, WN WA.

BUT, ON LONGER TIME SCALES, RESISTIVITY (AND OTHER NON-IDEAL EFFECTS) ARE IMPORTANT.

 $\nabla \mathbf{X} \mathbf{B} = \mu_0 \overline{\mathbf{J}}$ $p\frac{dv}{dt} = -\nabla P + 5 X R$ $\frac{2\overline{B}}{2t} = -\nabla \times \overline{E}$ $\overline{E} = -\nabla \times \overline{B} + \overline{D}\overline{J}$ WHAT IS ?? FOR A CONSTANT E, COLLISIONS LEAD TO A CONSTANT DRIFT O. 7 ml. m VELOCITY. $\begin{pmatrix} \mathcal{M} = \frac{me}{e^2 \pi T_e} \times \frac{1}{T_e^{3/2}} \end{pmatrix} \mathcal{O} \approx g \overline{E} + m \overline{V}_{teol} =) \overline{V} = \frac{g \overline{T_{cut}} \overline{E}}{m} \overline{E}$ $\mathcal{U} = \frac{g \overline{T_{cut}} \overline{E}}{m} \overline{E}$ $\mathcal{U} = \frac{g \overline{T_{cut}} \overline{E}}{m} \overline{E}$ $\mathcal{U} = \frac{g \overline{T_{cut}} \overline{E}}{m} \overline{E}$ 15



There is no "equilibrium" in resistive MHD MHOI

 $M_0 \nabla P = (\nabla \times B) \times \overline{D}$

Fonce BALAnce IS FIND EAST TO FIND EAST TO FIND i.o. GRAD-SHAFAANOU

FINITE RESISTIVITY IMPLIES NO EQUILIBRIUM PLASMA DISCHANGES (LINE) A TURAMAR MUST RE A TURAMAR MUST RE ACTIVELY SUCTIAINED PNOT SO FOR SOND STELLARATINS AND THER LEVITATED DIPSUR







"Simple" Resistive Equilibrium \Rightarrow Transport $\frac{\partial p}{\partial r} = -\nabla \cdot p \nabla$ $\begin{cases} \frac{\partial n}{\partial t} = \nabla \times (\nabla \times s) - \nabla \times (\frac{n}{ds} \nabla \times d) \end{cases}$



 $\frac{\overline{F}}{\overline{P}} = \frac{\overline{F}}{\overline{P}} + \frac{\overline{F}}{\overline{P}} = \frac{\overline{F}}{\overline{P}} + \frac{\overline{F}}{\overline{P}} = \frac{\overline{F}}{\overline{P}$ MOINY RADIAL POLOIDAL $\frac{\partial \rho}{\partial t} = -\nabla \left(\frac{\rho \pi}{R^2} D \times J \right)$ $= \nabla \left(\frac{\rho \pi}{R^2} D \times J \right)$ $= \nabla \left(\frac{\rho \pi}{R^2} \nabla_{+} \rho \right)$ $= \nabla \left(\frac{\rho \pi}{R^2} \nabla_{+} \rho \right)$ $= \rho k T$ $= \left(\frac{h}{h_0} \right) D^2 B$ $= \frac{h}{R_0} \left[\nabla \left(\sigma \cdot g \right) - \nabla^2 g \right] - \nabla \left(\frac{h}{k_0} \right) \times J$ $= \left(\frac{h}{h_0} \right) D^2 B$ $= \frac{h}{R_0} D^2 B$ $\frac{1}{1} = \frac{pn}{p^2} \sqrt{\frac{pn}{m_c}} = \frac{m}{\mu_0} \left(\frac{\sqrt{m_c}}{\sqrt{\frac{p}{A}}} \right)$ DXXDm Dr XDm Dr Dm P a (M/no) A

"Simple" Resistive Equilibrium \Rightarrow Transport Dm n Mo -> CURRENT RELAXES ON THE TIME SCALE OF TM 2/20m DEVICE SIZE, (27) T_e , (Kev) $m(m^{-3}) h T_m$ (Sec) 0.05 1×10 3×10 8msEC 0.14 19 3×10 8 600, SEC HBT-EP

0.13 T-3

0.85 DIII-D

JET 1.6

2.6 FTER

77 (coppin) = 1.7 × 10 7 m





₩6

15

10

19 2×10 9 4005EC

1 × 10²⁰ 6× 10^{1.5} Hours 1 × 10²⁰ 1× 10⁹ 2.2 Hours



Tearing Modes: Internal Resonances ($B \cdot \nabla = 0$)

REVIEW

NOUCTION $\frac{2\overline{B}}{\partial t} = -\nabla \times E \longrightarrow \frac{2\overline{B}}{2t}$ do

FOR EXTERNAL MODES B. 7 = O ANYWHERE I

REDUCED MHD

FOR INTERNAL MODES

B.D = O INSIDE PL



$$= \nabla \times (\nabla \times B) - \nabla \times (\pi \overline{J})$$
$$= (B \cdot \overline{B}) \overline{\nabla} - \nabla \times (\pi \overline{J})$$
$$1 = (\overline{B} \cdot \overline{B})$$



INTERNAL



Magnetic Islands



Figure 9.5 Equally-spaced contours of the magnetic flux-function, $\hat{\psi}(\hat{x},\xi)$, in the vicinity of a constant- ψ magnetic island chain. The magnetic separatrix is shown as a dashed line. 21

Sec. 9.7 Fitzpatrick

Do Magnetic Islands Grow? Or Shrink?



Do Magnetic Islands Grow? Or Shrink?

 Δ ' determines structure of surface current, *K*, at resonant surface.

rs.

How to calculate $\Delta'?$

LINEAR RESPONSE:

How to calculate Δ ?

INNER N PLASMA CURRENT PRADIENT DRIVES TEARING MODES

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Tearing Mode Dynamics in Tokamak Plasmas

Author Richard Fitzpatrick

Published June 2023

Download ebook

The development of humankind's ultimate energy source, nuclear fusion, has proceeded slowly but surely over the course of the last 60 years. This comprehensive book aims to outline a realistic, comprehensive, self-consistent, analytic theory of tearing mode dynamics in tokamak plasmas. It discusses a fluid theory of a highly magnetized plasma that treats the electrons and ions as independent fluids, and then proceeds to develop the theory of tearing modes, first approximating the geometry of a tokamak plasma as a periodic cylinder, but eventually considering the toroidal structure of real tokamak plasmas. This book also describes the stability of tearing modes, the saturation of such modes, and the evolution of their phase velocity due to interaction with other tearing modes, as well as the resistive vacuum vessel, and imperfections in the tokamak's magnetic field. This text will appeal to scientists and graduate students engaged in nuclear fusion research, and would make a useful reference for graduate plasma physics courses. Part of IOP Series in Plasma Physics.

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Topics Not Covered...

- Kinetic theory, Vlasov equation, wave-particle interactions, quasilinear theory, particle trapping, ...
- Gyrokinetics, drift waves, drift-wave transport, ...
- Nonlinear topics, like shocks, wave-echoes, reconnection, energy-momentum transfer, turbulence, ...
- Computational plasma physics, modeling, validation and verification, ...
- Relativistic physics, intense laser-plasma, astrophysical objects, particle acceleration, ...
- Plasma applications, like surface processing, chemistry, medicine, lighting, electric propulsion. ...
- Fusion energy challenges, like plasma-wall interactions, fueling, heating, control, ...
- Instrumentation, diagnostics, plasma reactors, ...

Next Lecture: Final Exam Review